Neutron-antineutron oscillations in the deuteron studied with NN and $\bar{N}N$ interactions based on chiral effective field theory*

Johann Haidenbauer^{1;1)} Ulf-G. Meißner^{2,1,3}

¹Institute for Advanced Simulation, Institut für Kernphysik and Jülich Center for Hadron Physics,
Forschungszentrum Jülich, D-52425 Jülich, Germany

²Helmholtz-Institut für Strahlen- und Kernphysik and Bethe Center for Theoretical Physics, Universität Bonn, 53115 Bonn, Germany

³Tbilisi State University, 0186 Tbilisi, Georgia

Abstract: Neutron-antineutron $(n-\bar{n})$ oscillations in the deuteron are considered. Specifically, the deuteron lifetime is calculated in terms of the free-space $n-\bar{n}$ oscillation time $\tau_{n-\bar{n}}$ based on NN and $\bar{N}N$ interactions derived within chiral effective field theory (EFT). This results in $(2.6\pm0.1)\times10^{22}\,\tau_{n-\bar{n}}^2$ s, which is close to the value obtained by Dover and collaborators more than three decades ago, but disagrees with recent EFT calculations that were performed within the perturbative scheme proposed by Kaplan, Savage, and Wise. Possible reasons for the difference are discussed.

Keywords: neutron-antineutron oscillations, chiral effective field theory, NN interaction, antinucleon-nucleon interaction

DOI: 10.1088/1674-1137/44/3/033101

1 Introduction

Neutron-antineutron $(n-\bar{n})$ oscillations involve a change of the baryon number (B) by two units $(|\Delta B| = 2)$. An experimental observation would allow a glimpse on physics beyond the standard model, see e.g. [1]. Since in such oscillations B is violated the process satisfies one of the Sakharov conditions [2] that have been formulated in order to explain the observation that there is more matter than anti-matter in the universe. Given such important and far-reaching consequences it is not surprising that there is a vast amount of literature on this topic [3]. Indeed, even within the past two years a wealth of papers have been published that deal with various and quite different aspects of $n-\bar{n}$ oscillations [4-11].

The key quanitity in this subject is the free $n - \bar{n}$ oscillation time, $\tau_{n-\bar{n}}$. The presently best experimental limit on it is $\tau_{n-\bar{n}} > 0.86 \times 10^8$ s ≈ 2.7 yr (with 90% C.L.) [12]. Additional information can be gained by studying $n - \bar{n}$ oscillations in a nuclear environment. Corresponding experiments have been performed, e.g., for 56 Fe [13], 16 O [14],

and for the deuteron (2 H) [15], while others are planned [16]. In such a case the oscillation process is suppressed as compared to the free situation. The pertinent lifetime τ_{nuc} is commonly expressed in terms of the one in free space as [3]

$$\tau_{\text{nuc}} = R\tau_{n-\bar{n}}^2 \,, \tag{1}$$

where R is an intranuclear suppression factor, also called reduced lifetime, that depends on the specific nucleus. It can be calculated from nuclear theory and then can be used to relate the measured lifetimes of those nuclei with the free $n - \bar{n}$ oscillation time [3], see, e.g., Refs. [17-21].

For a long time the suppression factors published in Ref. [18] have been used as standard by experimentalists in the interpretation of their measurements [13,15]. For example, in case of the deuteron the corresponding value is $R \sim (2.40 - 2.56) \times 10^{22} \text{ s}^{-1}$, a prediction based on the phenomenological antinucleon-nucleon ($\bar{N}N$) potentials by Dover and Richard [22,23]. Recently, however, those values have been called into question in a work by Oosterhof et al. [7]. In that study an effective field theory for the $|\Delta B| = 2$ interaction is constructed and the quantity R

Received 6 November 2019, Published online 7 January 2020

^{*} This work is supported in part by the DFG and the NSFC through funds provided to the Sino-German CRC 110 "Symmetries and the Emergence of Structure in QCD" (DFG grant. no. TRR 110) and the VolkswagenStiftung (93562). The work of UGM was supported in part by The Chinese Academy of Sciences (CAS) President's International Fellowship Initiative (PIFI) (2018DM0034)

¹⁾ E-mail: j.haidenbauer@fz-juelich.de

Content from this work may be used under the terms of the Creative Commons Attribution 3.0 licence. Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI. Article funded by SCOAP3 and published under licence by Chinese Physical Society and the Institute of High Energy Physics of the Chinese Academy of Sciences and the Institute of Modern Physics of the Chinese Academy of Sciences and IOP Publishing Ltd

is evaluated within the power counting scheme proposed by Kaplan, Savage, and Wise (KSW) [24,25] for the nucleon-nucleon (NN) and $\bar{N}N$ interactions. The value of R for the deuteron obtained in that approach is with $(1.1 \pm 0.3) \times 10^{22} \text{ s}^{-1}$ about a factor 2 smaller than the one by Dover et al. [18].

In the present paper we report on a calculation of the deuteron annihilation lifetime, considering neutron-antineutron oscillations. The main motivation comes, of course, from the aforementioned discrepancy and the work aims at providing if possible a plausible explanation for the difference. In our study we follow closely the benchmark calculations of Sandars [17] and Dover et al. [18]. The essential new aspect is that we employ modern interactions for the involved NN and the $\bar{N}N$ systems. Modern means that these interactions have been derived in a consistent and systematic framework, namely chiral effective field theory (EFT) [26]. And it means that these potentials are in line with present-day empirical information. This concerns specifically the $\bar{N}N$ system where a wealth of data on $\bar{p}p$ scattering and the charge-exchange reaction $\bar{p}p \rightarrow \bar{n}n$ has accumulated in the years after the publication of Ref. [18], notably due to measurements at the LEAR facility at CERN [27]. That fact is accounted for by utilizing $\bar{N}N$ potentials which have been fitted to up-to-date phase shifts and inelasticities provided by a recently published phase-shift analysis of available $\bar{p}p$ scattering data [28].

The $\bar{N}N$ interactions used in the present study are taken from two works [29,30]. In the first reference, $\bar{N}N$ potentials up to next-to-next-to-leading order (N²LO) were constructed, based on a modified Weinberg power counting, in close analogy to pertinent studies of the NN interaction [31]. In the second, the study was extended to next-to-next-to-leading order (N³LO) and, in addition, a new regularization scheme was implemented that had been introduced in the NN study of Ref. [32]. In the actual calculations we rely mostly on the recent more refined potential of higher order [30] that describes $\bar{N}N$ phase shifts and scattering observables up to laboratory energies of $T_{\rm lab} \sim 250$ MeV. However, additional calculations with the other potential are performed in order to estimate possible uncertainties that arise, e.g., from the employed regularization scheme.

The paper is structured in the following way: In Sec. 2 a basic description of the employed formalism is provided. Our results are presented and discussed in Sec. 3. We compare also with the works of [18] and [7]. The paper closes with a brief summary.

2 Formalism

In our treatment of $n-\bar{n}$ oscillations in the deuteron we follow very closely the formalism outlined in Refs.

[17,18]. However, contrary to those works our calculation is performed in momentum space and, therefore, we provide details about the main steps below. The starting point is the eigenvalue (Schrödinger) equation [17]

$$\begin{pmatrix} H_0 + V_{np} & V_{n-\bar{n}} \\ V_{n-\bar{n}} & H_0 + V_{\bar{n}p} \end{pmatrix} \begin{pmatrix} |\psi_{np}\rangle \\ |\psi_{\bar{n}p}\rangle \end{pmatrix} = (E - i\Gamma/2) \begin{pmatrix} |\psi_{np}\rangle \\ |\psi_{\bar{n}p}\rangle \end{pmatrix}.$$
(2)

Here, V_{np} and $V_{\bar{n}p}$ are the potentials in the np and $\bar{n}p$ systems and $|\psi_{np}\rangle$ and $|\psi_{\bar{n}p}\rangle$ are the corresponding wave functions. The systems are coupled via $V_{n-\bar{n}}$ which is given by the $n-\bar{n}$ transition matrix element $\delta m_{n-\bar{n}}$ where the latter is proportional to the inverse of the $n-\bar{n}$ oscillation time, i.e. $V_{n-\bar{n}} = \delta m_{n-\bar{n}} = \hbar/\tau_{n-\bar{n}}$ [3].

To leading order the $\bar{n}p$ component $|\psi_{\bar{n}p}\rangle$ obeys the equation

$$(H_0 + V_{\bar{n}p} - E_d)|\psi_{\bar{n}p}\rangle = -V_{n-\bar{n}}|\psi_d\rangle, \qquad (3)$$

where E_d is the unperturbed energy of the deuteron and $|\psi_d\rangle$ is the corresponding deuteron wave function. The decay width of the deuteron, Γ_d , is then [17]

$$\Gamma_d = -2 V_{n-\bar{n}} \operatorname{Im} \langle \psi_d | \psi_{\bar{n}p} \rangle . \tag{4}$$

We solve Eq. (3) in momentum space. Performing a partial wave decomposition and taking into account the coupling of the 3S_1 and 3D_1 channels, the above integral equation reads

$$(2E_{p} - E_{d})\psi_{\bar{n}p}^{L}(p) + \sum_{L'} \int \frac{\mathrm{d}q \, q^{2}}{(2\pi)^{3}} V_{\bar{n}p}^{L,L'}(p,q)\psi_{\bar{n}p}^{L'}(q)$$

$$= -V_{n-\bar{n}}\psi_{d}^{L}(p) , \qquad (5)$$

with L, L' = 0, 2. Note that E_d is the total energy corresponding to the deuteron, i.e. $E_d - 2m_N = 2\sqrt{m_N^2 - \kappa^2} - 2m_N = -B_d$ where B_d is the standard binding energy of 2.225 MeV and $\kappa = \sqrt{m_N B_d} \simeq 45.7$ MeV is the binding momentum. The deuteron wave function is normalized by

$$\int dp p^2 \left[(\psi_d^0(p))^2 + \psi_d^2(p))^2 \right] = 1$$
 (6)

and the width is provided by

$$\Gamma_d = -2 V_{n-\bar{n}} \operatorname{Im} \sum_L \int dp p^2 \psi_d^L(p) \psi_{\bar{n}p}^L(p) .$$
 (7)

The deuteron lifetime τ_d is given by $\tau_d = \hbar/\Gamma_d$. The interesting quantity is the so-called reduced lifetime R [17,18,21] which relates the free $n-\bar{n}$ oscillation lifetime with that of the deuteron,

$$\tau_d = R \tau_{n-\bar{n}}^2$$
 i.e. $R = \frac{\hbar}{\Gamma_d \tau_{n-\bar{n}}^2}$ (8)

3 Results and discussion

As already stated in the Introduction, we use the $\bar{N}N$ interactions from Refs. [29,30] derived within chiral EFT

and the deuteron wave functions from the corresponding NN potentials [31,32], derived in the same framework. We consider the two $\bar{N}N$ interactions because they are based on rather different regularization schemes. In the earlier potential [29] a non-local exponential exponential regulator was employed for the whole potential while in the recent work [30] a local regulator was adopted for the evaluation of the one- and two-pion contributions, see Refs. [30,32] for details. Comparing the pertinent results will allow us to shed light on the question in how far the choice of the regulator influences the predictions. For exploring further the sensitivity of the results to the deuteron wave function we employ also those of two meson-exchange potentials [33,34].

Our results are summarized in Table 1. They are based on our N³LO interaction with cutoff $R_0 = 0.9$ fm from Ref. [30] and the N²LO interaction with cutoff $\{\Lambda, \tilde{\Lambda}\} = \{450, 500\}$ MeV from Ref. [29]. For details on those interactions we refer the reader to the corresponding publications. Besides the predictions for R based on the chiral $\bar{N}N$ interactions we list also the values given in Ref. [18] where the $\bar{N}N$ potentials DR₁ and DR₂ by Dover-Richard [22,23] have been utilized. Furthermore we include results from the calculation of Oosterhof et al. performed directly within EFT on the basis of the KSW approach. In this case R can be represented in a compact analytical form which reads up to NLO [7]

$$R = -\frac{\kappa}{m_N} \frac{1}{\text{Im} \, a_{\bar{n}p}} \frac{1}{1 + 0.4 + 2\kappa \,\text{Re} \, a_{\bar{n}p} - 0.13 \pm 0.4} \,. \tag{9}$$

Obviously, the only parameter here is the $\bar{n}p$ 3S_1 scattering length. All other quantities that enter are well established *NN* observables, cf. Ref. [7] for details. Note that in that paper, the scattering length Re $a_{\bar{n}p}$ was taken from Ref. [30].

As can be seen from Table 1 the values for R predicted by the chiral $\bar{N}N$ interactions are fairly similar to those obtained for the DR potentials in the past. The results based on the framework employed by Oosterhof et al. [7], utilizing the scattering length from the N³LO chiral $\bar{N}N$ interaction [30], on the other hand, are rather different. This suggests that the large discrepancy observed in Ref. [7] is certainly not due to differences in Im $a_{\bar{n}p}$ but primarily a consequence of the different approaches.

For illustration we present in Fig. 1 the wave functions of the ${}^{3}S_{1}$ component of the deuteron and of the cor-

responding $\bar{n}p$ state with arbitrary normalizations. In case of the latter the imaginary part is shown which is relevant for the determination of the width, cf. Eq. (7). The impact of the different regularization schemes used in Refs. [29,31] and [30,32], respectively, can be clearly seen from the different behavior of the wave functions for large momenta. Note, however, that the bulk contribution to the integral in Eq. (7) comes from momenta $p \le 300$ MeV/c. Contributions from larger momenta to Γ_d are only in the order of 5% as revealed by test calculations.

In order to investigate the sensitivity of our results to the utilized input we performed various exploratory calculations. First of all, we employed the NLO and N²LO variants of the considered $\bar{N}N$ (and NN) interactions. The corresponding predictions for R were found to lie within a range of $(2.48-2.65)\times10^{22}$ s⁻¹. Taking this variation as measure for the uncertainty due to the nuclear structure, i.e. the NN and $\bar{N}N$ interactions (wave functions), leads to a value of roughly $R = (2.6 \pm 0.1) \times 10^{22} \text{ s}^{-1}$. Applying the method proposed in Ref. [32] for estimating the uncertainty to the calculation based on the $\bar{N}N$ interaction from 2017 [30], say, would actually lead to a slightly smaller uncertainty. We have also varied the deuteron wave functions alone. As an extreme case we even took wave functions from phenomenological NN potentials derived in an entirely different framework, namely in the meson-exchange picture [33,34]. Also here the obtained values for R stayed within the range given above. Finally, omitting the D-wave component of the deuteron wave function, which is kept in our calculation and in the one by Dover et al. [18] causes a 5% variation. But it leads to an increase of the value of R and, thus, does not bring it closer to the values presented by Oosterhof et al. Overall, we confirm the observation by Dover et al. that the predictions for R are fairly insensitive to the details of the employed $\bar{N}N$ potentials [18], provided that these potentials describe the $\bar{p}p$ data at low energies.

Since in Ref. [7] the $\bar{N}N$ interaction is represented by the leading-order term, i.e. the scattering length alone, we have evaluated the effective range parameter for our $\bar{N}N$ interactions and used it to extrapolate the $\bar{N}N$ amplitude to the deuteron pole. The found variations are in the order of 10 % and can certainly not explain the large difference documented in Table 1.

We do not have a clear explanation for the difference

Table 1. Reduced lifetime R calculated for the χ EFT $\bar{N}N$ potentials from Refs. [29,30], together with information on the pertinent $\bar{n}p$ 3S_1 scattering length. Results for the Dover-Richard potentials DR₁ and DR₂ are taken from Ref. [18]. The corresponding scattering lengths are from Ref. [35]. Predictions based on Eq. (9), i.e. on the KSW approach applied in Ref. [7], are indicated too.

	χ EFT N ² LO [29]	χ EFT N ³ LO [30]	DR ₁ [18]	DR ₂ [18]
$R(s^{-1})$	2.49×10^{22}	2.56×10^{22}	2.56×10^{22}	2.40×10^{22}
(Eq. (9))	$(1.1 \pm 0.3) \times 10^{22}$	$(1.2 \pm 0.3) \times 10^{22}$	$(1.4 \pm 0.4) \times 10^{22}$	$(1.3 \pm 0.3) \times 10^{22}$
$a_{^3S_1}(\text{fm})$	0.44 – i 0.91	0.44-i0.96	0.87 - i 0.66	0.89 - i 0.71

of our results (and those of Ref. [18]) to the ones of Oosterhof et al. [7]. However, we believe that it is due to the fact that in the latter work the width Γ_d is evaluated following the perturbative scheme developed by Kaplan, Savage, and Wise [25]. In that scheme there is no proper deuteron wave function. Rather one works with an effectively constructed wave function that is represented in terms of an irreducible two-point function [7,25]. This seems to work well for some electromagnetic form factors of the deuteron, at least at low momentum transfer [25,36]. On the other hand, the quadrupole moment of the deuteron is overestimated by 40% [25], which hints that the properties of the wave function at large distances (small momenta) are not that well represented in this scheme. Clearly, this should have an impact on the quantity studied in the present work as well. Note that a comparable agreement (mismatch) with regard to the KSW scheme has been also observed in studies of the electric dipole moment (magnetic quadrupole moment) of the deuteron [37-39]. In any case, one should not forget that there is convergence problem of the KSW approach for NN partial waves where the tensor force from pion exchange is present [40]. It affects specifically the ${}^{3}S_{1}$ - ${}^{3}D_{1}$ channel where difficulties appear already for momenta around 100 MeV/c, see [40] and also the discussions in Refs. [26,41]. Interestingly enough, this is the momentum region where the dominant contributions in our calculation come from, see Fig. 1.

In contrast to [7] in our study the $\bar{N}N$ (NN) and $|\Delta B| = 2$ interactions are not treated at the same order. Guided by the success of the work by Epelbaum et al. [31,32] based on the Weinberg counting the non-perturbative effects due to the NN and $\bar{N}N$ interactions are fully take into account from the very beginning. It is worth mentioning that in this approach the quadrupole moment of the deuteron is very close to its empirical value already for interactions at NLO [31,32].

Finally, note that in the present work we have only considered the contribution from actual $n - \bar{n}$ oscillations to the deuteron decay rate. In principle, there can be also contributions from a direct annihilation of the NN system. Moreover, there can be a $NN \rightarrow \bar{N}N$ transition involving both nucleons and the $\bar{N}N$ state. The latter gives rise to a $|\Delta B| = 2$ four-baryon contact term involving an unknown complex low-energy constant, which has been discussed and considered in the work by Oosterhof et al. [7]. We refrain from introducing such a contribution in the present study which, anyway, is of higher order in the employed counting scheme. In this context, we would like to mention that for most $|\Delta B| = 2$ sources considered in Ref. [7] the deuteron decay rate was found to be dominated by free $n-\bar{n}$ oscillations. Nonetheless, an uncertainty in the order of $\pm 0.3 \times 10^{22} \text{ s}^{-1}$ for the reduced lifetime *R* due to additional $|\Delta B| = 2$ mechanisms, as suggested in that work, has to be certainly expected. Note that additional

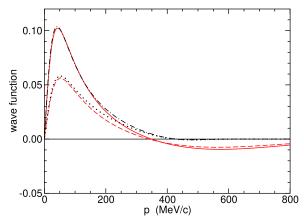


Fig. 1. (color online) Wave functions of the deuteron (ψ_d) ; upper curves) and of the $\bar{n}p$ state in the 3S_1 partial wave (imaginary part of $\psi_{\bar{n}p}$; lower curves), multiplied by the momentum p and with arbitrary normalization. Solid and dashed (red) curves are for the N³LO $\bar{N}N$ interaction [30] with cutoff $R_0 = 0.9$ fm and dash-dotted and dotted (black) curves for the N²LO $\bar{N}N$ interacton [29] with cutoff $\{\Lambda, \tilde{\Lambda}\} = \{450, 500\}$ MeV.

decay mechanisms should always increase the decay width of the deuteron and, thus, lead to smaller values of the reduced lifetime R.

4 Summary

In the present work we have considered neutron-antineutron oscillations in the deuteron. In particular, we have calculated the deuteron lifetime in terms of the free-space $n-\bar{n}$ oscillation time based on $\bar{N}N$ [29,30] and NN [31,32] interactions derived within chiral effective field theory. The value obtained for the so-called reduced lifetime R which relates the free-space $n-\bar{n}$ oscillation time $\tau_{n-\bar{n}}$ with the deuteron lifetime is found to be $R = (2.6 \pm 0.1) \times 10^{22} \text{ s}^{-1}$, where the quoted uncertainty is due to the NN and $\bar{N}N$ interactions (wave functions).

Our result for *R* is close to the value obtained by Dover and collaborators more than three decades ago [18] but disagrees with recent EFT calculations, performed within the perturbative scheme proposed by Kaplan, Savage, and Wise [7], by about a factor of 2. We believe that the difference is due to the fact that in the perturbative KSW scheme there is no proper deuteron wave function. Rather this ingredient is represented effectively in terms of an irreducible two-point function. It is known from past studies that the KSW approach fails to describe quantities that depend more sensitively on the wave function like, for example, the quadrupole moment of the deuteron [25].

We acknowledge stimulating discussions with Jordy de Vries, Andreas Nogga and Tom Luu.

References

- R. N. Mohapatra, J. Phys. G, 36: 104006 (2009)
- A.D. Sakharov, JETP Lett., 5: 24 (1967)
- D. G. Phillips II et al, Phys. Rept., 612: 1 (2016)
- J. Bijnens and E. Kofoed, Eur. Phys. J. C, 77: 867 (2017) S. Gardner and X. Yan, Phys. Rev. D, 97: 056008 (2018)
- C. Grojean, B. Shakya, J. D. Wells et al, Phys. Rev. Lett., 121: 171801 (2018)
- F. Oosterhof, B. Long, J. de Vries et al, Phys. Rev. Lett., 122: 172501 (2019)
- E. Rinaldi, S. Syritsyn, M. L. Wagman et al, Phys. Rev. D, 99: 074510 (2019)
- S. Gardner and X. Yan, Phys. Lett. B, 790: 421 (2019)
- Z. Berezhiani and A. Vainshtein, Phys. Lett. B, 788: 58 (2019) 10
- B. O. Kerbikov, Phys. Lett. B, 795: 362 (2019)
- M. Baldo-Ceolin et al, Z. Phys. C, 63: 409 (1994)
- J. Chung et al, Phys. Rev. D, 66: 032004 (2002)
- K. Abe et al (Super-Kamiokande Collaboration), Phys. Rev. D, 91: 072006 (2015)
- B. Aharmim et al (SNO Collaboration), Phys. Rev. D, 96: 092005 (2017)
- Barrow, E. S. Golubeva, E. Paryev et al, arXiv: J. L. 1906.02833[hep-ex]
- P. G. H. Sandars, J. Phys. G, 6: L161 (1980) 17
- C. B. Dover, A. Gal, and J.-M. Richard, Phys. Rev. D, 27: 1090
- 19 W. M. Alberico, A. De Pace, and M. Pignone, Nucl. Phys. A, 523: 488 (1991)
- J. Hufner and B. Z. Kopeliovich, Mod. Phys. Lett. A, 13: 2385 (1998)
- E. Friedman and A. Gal, Phys. Rev. D, 78: 016002 (2008)
- C. B. Dover and J.-M. Richard, Phys. Rev. C, 21: 1466 (1980)
- C. B. Dover and J.-M. Richard, *Phys. Rev. C*, **25**: 1952 (1982)

- D. B. Kaplan, M. J. Savage, and M. B. Wise, Phys. Lett. B, 424: 390 (1998)
- 25 D. B. Kaplan, M. J. Savage, and M. B. Wise, Phys. Rev. C, 59: 617 (1999)
- E. Epelbaum, H. W. Hammer, and U.-G. Meißner, Rev. Mod. 26 Phys., 81: 1773 (2009)
- E. Klempt, F. Bradamante, A. Martin et al, Phys. Rept., 368: 119 (2002)
- D. Zhou and R. G. E. Timmermans, Phys. Rev. C, 86: 044003 (2012)
- 29 X. W. Kang, J. Haidenbauer, and U.-G. Meißner, JHEP, 1402: 113 (2014)
- L. Y. Dai, J. Haidenbauer, and U.-G. Meißner, JHEP, 1707: 078 30 (2017)
- E. Epelbaum, W. Glöckle, and U.-G. Meißner, Nucl. Phys. A, 747: 362 (2005)
- E. Epelbaum, H. Krebs, and U.-G. Meißner, Eur. Phys. J. A, 51: 53 (2015)
- J. Haidenbauer, K. Holinde, and M. B. Johnson, Phys. Rev. C, 48: 2190 (1993)
- R. Machleidt, Phys. Rev. C, 63: 024001 (2001)
- J. Carbonell, J.-M. Richard, and S. Wycech, Z. Phys. A, 343: 325
- 36 M. Walzl and U.-G. Meißner, Phys. Lett. B, 513: 37 (2001)
- J. de Vries, E. Mereghetti, R. G. E. Timmermans et al, Phys. Rev. Lett., **107**: 091804 (2011)
- C.-P. Liu, J. de Vries, E. Mereghetti et al, Phys. Lett. B, 713: 447 (2012)
- J. Bsaisou, J. de Vries, C. Hanhart et al, JHEP 1503: 104 (2015) 39 Erratum: [JHEP 1505, 083 (2015)]
- S. Fleming, T. Mehen, and I. W. Stewart, Nucl. Phys. A, 677: 313
- 41 P. F. Bedaque and U. van Kolck, Ann. Rev. Nucl. Part. Sci., 52: 339 (2002)